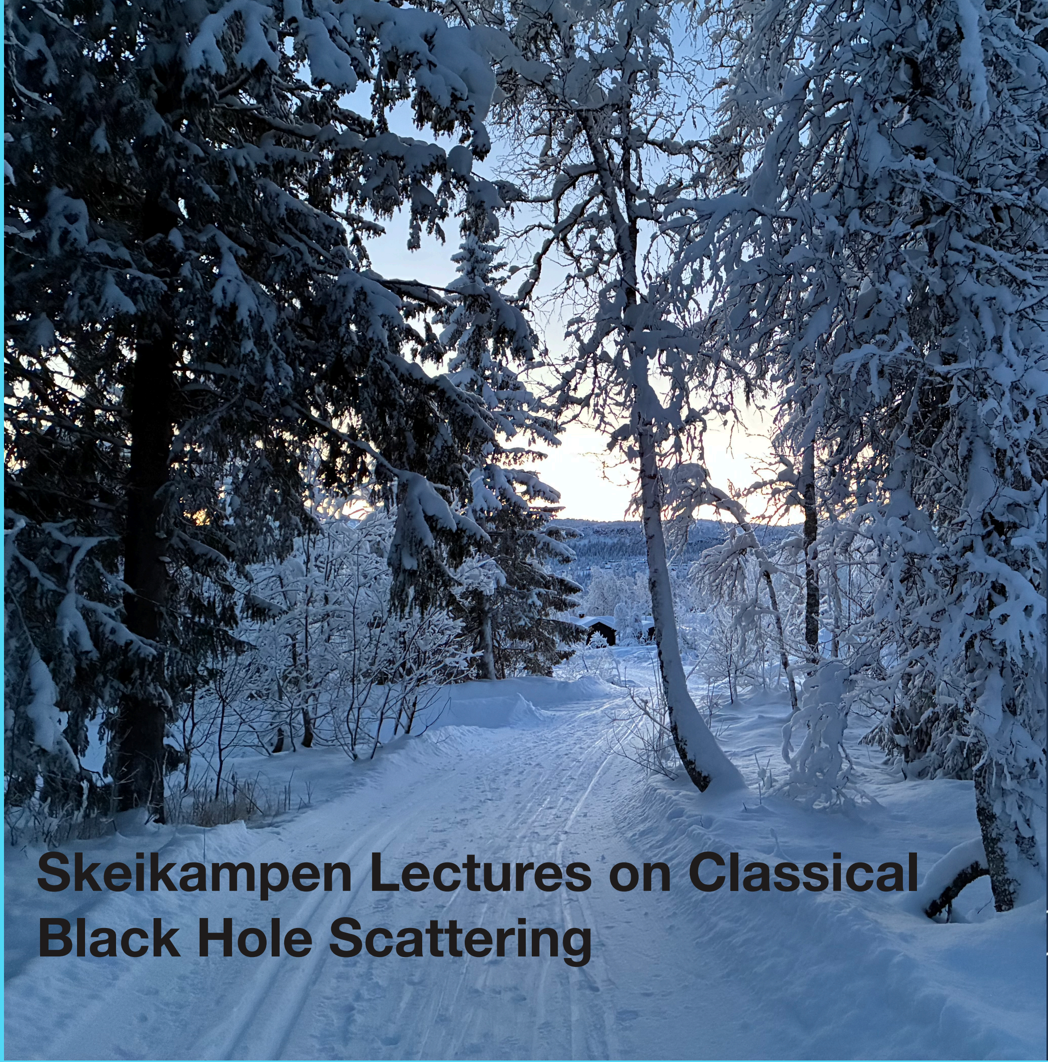
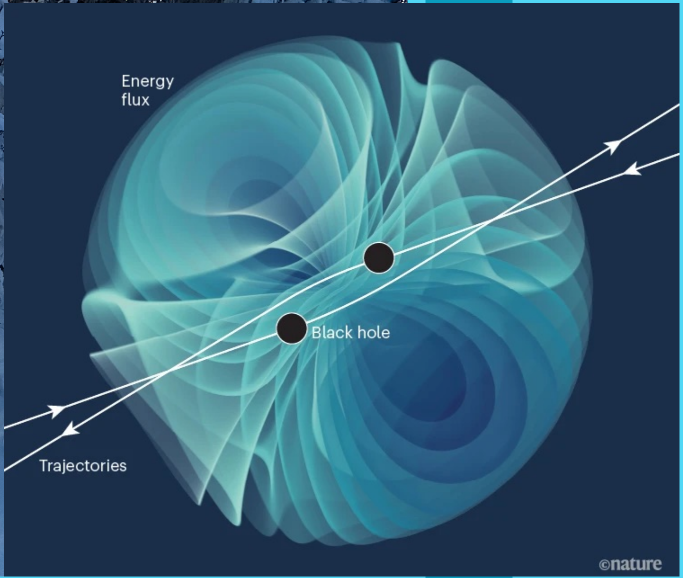


Skeikampen Lectures on Classical Black Hole Scattering



PLAN:

1) WORLDLINE EFT & GRAVITATIONAL 2-BODY PROBLEM

2) BH/NS SCATTERING: WORLDLINE QFT APPROACH:

[Diagrammar; Observables; GSF expansion; Feynman integral structure]

3) IMPULSE Δp @ 1PM & 2PM

4) HIGH PM WORKFLOW: INTEGRAND, IBPs, DE, BOUNDARIES

5) SPIN

6) GEODESIC MOTION FROM WQFT

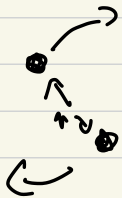
7) SCATTERING WAVEFORM @ LO.

REFERENCES: PhD thesis of GUSTAV UHRE JAKOBSEN 2308.04388

Physics Reports (2026?)

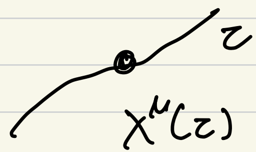
0) MOTIVATION → INTRO SLIDES.

1) WORLDLINE EFFECTIVE FIELD THEORY & THE GRAVITATIONAL 2-BODY PROBLEM



$Gm \ll c\, r$

Consider BHs/NSs as point particles



1) $$S_{BH} = -m \int ds = -m \int_{-\infty}^{\infty} d\tau \sqrt{\dot{x}^\mu \dot{x}^\nu g_{\mu\nu}}$$

Polyakov-trick:

$$\boxed{S_{BH} = \tilde{S}_{BH} = -\frac{m}{2} \int d\tau \left(e^{-1} \dot{x}^2 + e \right)}$$

$e(\tau)$ $$\frac{\delta \tilde{S}_{BH}}{\delta e} = 0 = 1 - e^{-2} \dot{x}^2 \Rightarrow \underline{e = \sqrt{\dot{x}^2}}$$

Reinsert e in $S_{BH}[e = \sqrt{\dot{x}^2}] = S_{BH}$

$\boxed{\text{Proper time gauge: } \quad e = 1}$ $\Rightarrow$ $\boxed{S_{BH} = -\frac{m}{2} \int d\tau \, \dot{x}^2}$

Constraint $\underline{1 = g_{\mu\nu} \, \dot{x}^\mu(\tau) \, \dot{x}^\nu(\tau)}$

Constraint is preserved under dynamics

$$\frac{d}{d\tau}\left(g_{\mu\nu} \, \dot{x}^\mu \dot{x}^\nu \right) = 0$$

E.o.m

<u>E.O.M:</u> $\boxed{\ddot{x}^\mu + \Gamma^\mu{}_{\beta\sigma} \, \dot{x}^\beta \dot{x}^\sigma = 0}$

2) $S_{GR} = \frac{1}{16\pi G} \int d^4x \, \sqrt{-g} \, R$ $\rightarrow$ EINSTEIN EQS.

$\Rightarrow$ $S = S_{BH1} + S_{BH2} + S_{GR} + S_{G.F.}$

WORLDLINE EFT DESCRIPTION OF BHs/NSs

$$S = -\frac{m}{2}\int d\tau\, g_{\mu\nu}\dot{x}^\mu \dot{x}^\nu + \left(\begin{matrix}\text{SPIN}\\ \text{DOFS}\end{matrix}\right) + \left(\begin{matrix}\text{TIDAL}\\ \text{EFFECTS}\end{matrix}\right) + \frac{1}{16\pi G_N}\int d^4x\sqrt{g}\,R$$

$$C_{E^2}\int d\tau \left(R_{\mu\dot{x}\nu\dot{x}}\right)^2 + C_{B^2}\left(\tilde{R}_{\mu\dot{x}\nu\dot{x}}\right)^2$$

SPINLESS SITUATION:

$$R_{\mu\nu\rho\sigma}\dot{x}^\nu = R_{\mu\dot{x}\rho\sigma}$$

3) WQFT DO EXPANSION IN "G"

$$\left(\frac{GM}{b}\right)^{\#}$$

Background field expansion:

$$x^\mu(\tau) = b^\mu + v^\mu\cdot\tau + z^\mu(\tau)$$

$$g_{\mu\nu}(x) = \eta_{\mu\nu} + \sqrt{G}\,h_{\mu\nu}(x)$$

Solve for z^μ & $h_{\mu\nu}$ by quantising them!

$$\langle \mathcal{O} \rangle_{\text{WQFT}} = \int \mathcal{D}[z^\mu, h_{\mu\nu}] \; \mathcal{O}[z,h] \; e^{-\frac{i}{\hbar} S[z,h]} \Big/ Z$$

$\Rightarrow$ TREE-LEVEL ONE POINT FUNCTIONS SOLVE

$\hbar \to 0$ CLASSICAL EQUATIONS OF MOTION

$$\langle z^\mu(\tau) \rangle = \sum_{n=1}^{\infty} G^n \, z^\mu_{(n)}(\tau)$$

CLASSICAL TRAJ.

$$\sqrt{G} \langle h_{\mu\nu}(x) \rangle = \sum_n G^n \, h_{(n)\mu\nu}$$

WAVEFORM

$$\langle x \rangle = \int dx \; x \; p(x)$$

WHY 1-POINT FUNCTIONS @ TREE LEVEL

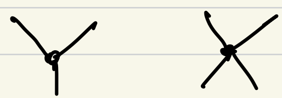
ϕ: Scalar Field

$$Z = \int \mathcal{D}\phi \; e^{-\frac{i}{\hbar} S[\phi]} = \int \mathcal{D}\phi \sum \frac{g^n \phi^n}{\hbar^n} \, e^{-\frac{i}{\hbar}\int \phi \partial^2 \phi}$$

$$S[\phi] = \int \frac{1}{2} \phi \partial^2 \phi + g_1 \phi^3 + g_2 \phi^4$$

$$\frac{1}{2}\int d^4x \, \partial_\mu \phi \, \partial^\mu \phi$$

$$\langle \phi \phi \rangle_{g=0} \sim \frac{\hbar}{p^2} \qquad x \overset{p}{\text{———}} y$$

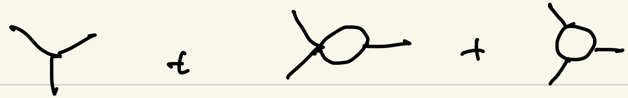


$\sim g_1/\hbar$ $\qquad$ $\sim g_2/\hbar$

Hence if we look at connected n-point fncts

$$\langle \phi_1 \dots \phi_n \rangle_{con} = \sum_{l=0}^{\infty} (\hbar)^{l+n-1} \; (\text{graphs with } n\text{-external legs and } n\text{-loops})$$

Makes sense only for $l=0$ & $n=1$ $\Rightarrow$ TREE-LEVEL 1-PT FCT.

<u>3-pt</u>: $\langle \phi\phi\phi\rangle =$ 

TREE 1-LOOP

$$\hbar^3 \frac{1}{\hbar} + \hbar^3 \frac{1}{\hbar^2}\cdot \hbar^2 \quad \hbar^3 \frac{1}{\hbar^3}\hbar^3 = \hbar^2 + \hbar^3$$

$$\boxed{\langle \phi_1 \dots \phi_N \rangle_{\text{DISCON}} \overset{\hbar\to 0}{=} \langle \phi_1\rangle^{\text{TREE}}_{\text{CON}} \langle \phi_2\rangle^{\text{TREE}}_{\text{CON}} \cdots \langle \phi_N\rangle^{\text{TREE}}_{\text{CON}}} + O(\hbar) + O(1/\hbar)$$

<u>SCHWINGER - DYSON - EQ!</u>

$$0 = \left\langle \frac{\delta S(\phi)}{\delta \phi} \right\rangle_{\hbar \to 0} \longrightarrow 0 = \frac{\delta S[\langle \phi \rangle^{\text{TREE}}]}{\delta \phi}$$

FEYNMAN DIAGRAMATICS

Worldline deflection:

$$S^{(0)}_{WL} = -\frac{m}{2}\int d\tau \left[\eta_{\mu\nu}\, \dot{x}^{\mu}(\tau)\, \dot{x}^{\nu}(\tau) + 1 \right]$$

$\hookrightarrow \dot{x}^{\mu}(\tau) = v^{\mu} + \dot{z}^{\mu}(\tau)$

$$= -\frac{m}{2}\int d\tau \left[1 + 2\, \dot{z}\cdot v + \dot{z}^2 \right]$$

$\hookrightarrow z(\infty) - z(-\infty) \to 0$

DROP $\Leftrightarrow$ (IN $\sim$ IN)

$$= -\frac{m}{2}\int d\tau \; \dot{z}^2$$

$$= -\frac{m}{2}\int_{\omega} \omega^2\, \tilde{z}(\omega)\cdot\tilde{z}(-\omega)$$

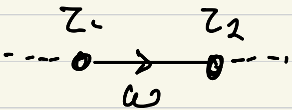
F.T:

$$z(\tau) = \int_{\omega} e^{-i\omega\cdot\tau}\, \tilde{z}(\omega)$$

$$\boxed{\left\langle z^{\mu}(\omega)\; z^{\nu}(-\omega) \right\rangle = -\frac{i}{m}\, \frac{\eta_{\mu\nu}}{(\omega + i0^{+})^2}}$$

PROPAGATOR

E.O.M: $\quad \underline{m\,\partial_\tau^2 X^\nu = \delta(\tau)} \quad \Leftrightarrow \quad i\, m\, \omega^2 X^\mu(\omega) = 1$



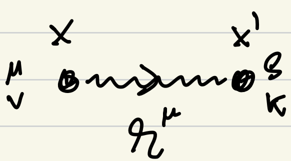
$$= -i\,\frac{\eta^{\mu\nu}}{m}\int_\omega \frac{e^{i\omega(\tau_1-\tau_2)}}{(\omega+i0^+)^2} = \frac{i\eta^{\mu\nu}}{2m}\Big(|\tau_1-\tau_2| + (\tau_1-\tau_2)\Big)$$

this is the $\wedge$ Green's function for $\quad m\,\partial_\tau^2 \quad$ on $\quad \mathbb{R}$

retarded

$\underline{\square h_{\mu\nu}(x) = 0}$

GRAVITON PROGATOR:



$$= \int_k \frac{i\,P_{\mu\nu\rho\kappa}\, e^{ik\cdot(x-x')}}{(k^0+i0^+)^2 - \vec{k}^2} \underset{D=4}{=} \frac{\delta[(x-x')^2]}{4\pi|\vec{x}-\vec{x}'|}\,P_{\mu\nu\rho\kappa}$$

$P^\rho P^\kappa \quad P_{\mu\nu\rho\kappa} = \frac{1}{2}\left(\eta_{\mu\rho}\eta_{\nu\kappa} + \eta_{\nu\rho}\eta_{\mu\kappa}\right) - \frac{1}{D-2}\,\eta_{\mu\nu}\eta_{\rho\kappa}$

VERTEX RULES:

(a) WORLDLINE VERTICES

$$\boxed{S_{WL}^{WT} = -\frac{m\kappa}{2}\int_\tau h_{\mu\nu}[x(\tau)]\,\dot{x}^\mu(\tau)\,\dot{x}^\nu(\tau)}$$

///: "quantum" fields

$$= \frac{m\kappa}{2}\int_\tau h_{\mu\nu}[x(\tau)]\left(v^\mu v^\nu + 2v^\mu \dot{z}^\nu(\tau) + \dot{z}^\mu(\tau)\,\dot{z}^\nu(\tau)\right)$$

$$\hookrightarrow\; h_{\mu\nu}[x(\tau)] = \int_k e^{ik\cdot(b+v\cdot\tau + z(\tau))}\,h_{\mu\nu}(-k)$$

$\int_k = \int \frac{dk}{(2\pi)}$ $\quad \hat{\delta}(x) = 2\pi\delta(x)$

$$z^\mu(\tau) = \int_\omega e^{i\omega\tau}\,z^\mu(\omega)$$

$$= \sum_{n=0}^{\infty}\frac{i^n}{n!}\int_k e^{ik(b+v\tau)}\left[k\cdot z(\tau)\right]^n h_{\mu\nu}(-k)$$

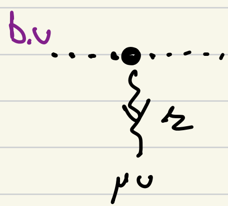
$\hookrightarrow$ F.T.

$$= \sum_{i=0}^{\infty} \frac{i^n}{n!} \int_{r,\omega_1 \ldots \omega_n} e^{ir\cdot b}\, e^{i\left(r\cdot v + \sum_i \omega_i\right)\cdot \tau} \left(\prod_{i=1}^{n} r\cdot z(-\omega_i)\right) h_{\mu\nu}(-r)$$

Insert this to get GRAVITON-WL VERTICES:

n=0:

$$S_{\text{INT}}\Big|_{z^0} = -\frac{m\kappa}{2} \int_r e^{ir\cdot b}\, \delta(r\cdot v)\, h_{\mu\nu}(-r)\, v^\mu v^\nu$$

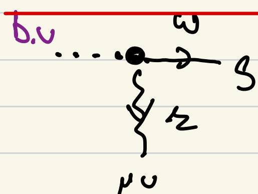


$$= -i\frac{m\kappa}{2}\, e^{ir\cdot b}\, \delta(r\cdot v)\, v^\mu v^\nu$$

$$\kappa^2 = 32\pi G$$

n=1:

$$S_{\text{INT}}\Big|_{z^1} = -i\frac{m\kappa}{2} \int_{r,\omega} e^{ir\cdot b}\, \delta(r\cdot v + \omega)\, h_{\mu\nu}(-r)\, z^\rho(-\omega) \left(2\omega\, v^{(\mu}\delta^{\nu)}_\rho + v^\mu v^\nu r_\rho\right)$$

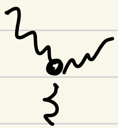
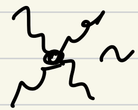
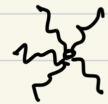


$$= \frac{m\kappa}{2}\, e^{ir\cdot b}\, \delta(r\cdot v + \omega) \left(2\omega\, v^{(\mu}\delta^{\nu)}_\rho + v^\mu v^\nu r_\rho\right)$$

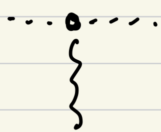
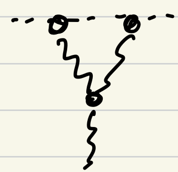
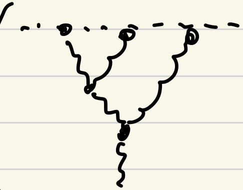
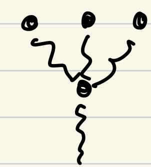
b) GRAVITON VERTICES

$$S_{EH} = \frac{1}{16\pi G}\int d^D x\, \sqrt{-g}\, R = \sum_{n=2}^{\infty} \sqrt{G}^{\,n-2}\ \text{"}\partial^2 h^n\text{"}$$

$g = \eta + \sqrt{G}\, h$

 $\sim \sqrt{G}$  $\sim G$  $\sim G^{3/2}$ ...

FUN WITH A SINGLE WORLDLINE

$$\sqrt{G}\,\langle h_{\mu\nu}\rangle =$$  $+$  $+ \Big($  $+$  $\Big) + \ldots$

$\mu\nu$ $\quad$ $\mu\nu$

G ✓ $\quad$ G^2 $\quad$ G^3

$g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}$

Yields PM expansion of Schwarzschild (harmonic coordinates)

AICHELBURG-SEXL FROM WQFT

Looks at massless point particle $\quad (e \to e \cdot m)$

$$S = -\frac{1}{2}\int d\tau \left[e^{-1}\,\dot{x}^2 - e\,\cancel{m}^2\right] \xrightarrow[m=0]{} -\frac{1}{2}\int d\tau\, e^{-1}\, g_{\mu\nu}\,\dot{x}^\mu \dot{x}^\nu$$

E.O.M for e: $\quad \boxed{0 = \dot{x}^2} \qquad e = 1/p \qquad x^\mu(\tau),\ e^{-1}(\tau)$

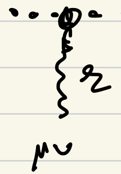
Energy of particle

$$S_{MASSLESS} = -\frac{p}{2}\int d\tau\; g_{\mu\nu}[x]\,\dot{x}^\mu \dot{x}^\nu$$

Background field exp. $\qquad x^\mu(\tau) = b^\mu + p\cdot\tau + z^\mu(\tau)$

$$p^2 = 0$$

$$\langle h_{\mu\nu}(x)\rangle = \int_r e^{ir\cdot x}$$



$$= \int_r e^{ir\cdot x}\left(-\frac{iK}{2}\,\delta(r\cdot P)\;P^\rho P^\kappa\right) i\,\frac{P_{\rho\kappa,\mu\nu}}{r^2}$$

$$= \frac{K}{2}\,P^\mu P^\nu \int_r e^{ir\cdot x}\,\delta(P\cdot r)\,\frac{1}{r^2}$$

Light-cone coord:

$$r^\mu = n^\mu(r\cdot\bar n) + \bar n^\mu(r\cdot n) + r_\perp^\mu \qquad P^\mu = P\,n^\mu$$

$$x^+ = n\cdot x \qquad x^- = \bar n\cdot x \qquad n\cdot\bar n = 1 \quad n^2 = 0 = \bar n^2$$

$$\int_r e^{ir\cdot x}\,\delta(P\cdot r)\,\frac{1}{r^2} = P^{-1}\int \frac{d^{D-2}r_\perp\,\cancel{dr}\,d\bar r}{(2\pi)^D}\, e^{i(\bar r\cdot x^- + \cancel{r\cdot x^+} - r_\perp\cdot x_\perp)}\,\frac{\delta(\cancel{r})}{2\cancel{r\bar r} - r_\perp^2}$$

$$= - \delta(x^-)\, p^{-1} \underbrace{\int \frac{d^{D-2}k_\perp}{(2\pi)^{D-2}} \frac{e^{-ik_\perp x_\perp}}{k_\perp^2}}$$

$D = 4 - 2\varepsilon$

$$= \frac{\Gamma(-\varepsilon)}{4\pi} \left(\pi x_\perp^2\right)^{\varepsilon} \underset{\varepsilon \to 0}{=} \log\left(x_\perp^2 / L^2\right)$$

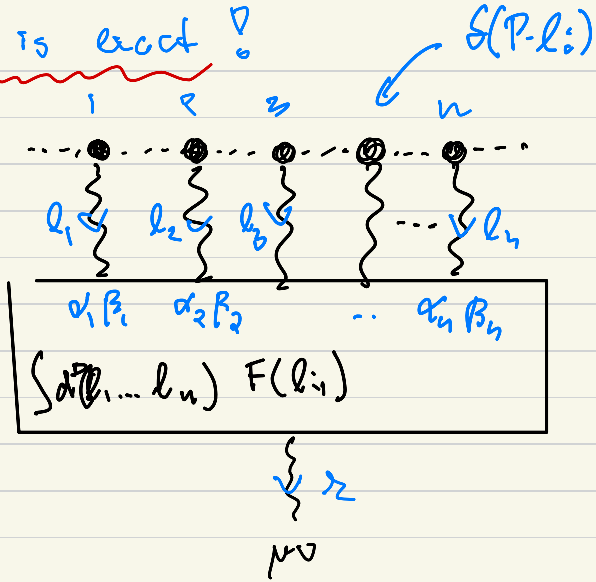
$$\Rightarrow \quad \kappa \langle h_{\mu\nu} \rangle = 4 G P^{-1} \delta(x^-) \log\left(\frac{x_\perp^2}{L^2}\right) p^\mu p^\nu$$

TAKE $\varepsilon \to 0$:

$$\boxed{ds^2 = 2 dx^+ dx^- - d\vec{x}^2 + 4 P G \delta(x^-) \log\left[\frac{\vec{x}^2}{4L^2}\right] (dx^-)^2}$$

AICHELBURG - SEXL SHOCKWAVE METRIC (1971)

This is exact !



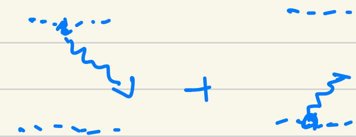
$$P^2 = 0 \qquad P\cdot l_i = 0$$

$$r = \sum_i l_i$$

$$\Omega_{\alpha_1\beta_1\dots\alpha_n\beta_n;\mu\nu}\, p^{\alpha_1} p^{\beta_1} \dots p^{\alpha_n} p^{\beta_n}$$

$$=$$

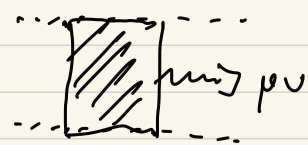
NOW THE 2-BODY PROBLEM

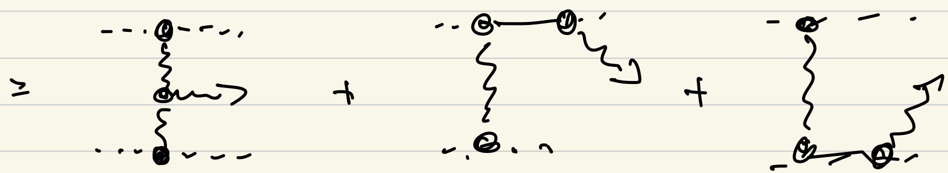


OBSERVABLES:

I WAVEFORM

$$\langle h_{\mu\nu}\rangle^{\text{TREE}} =$$





$G^{3/2}$

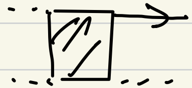
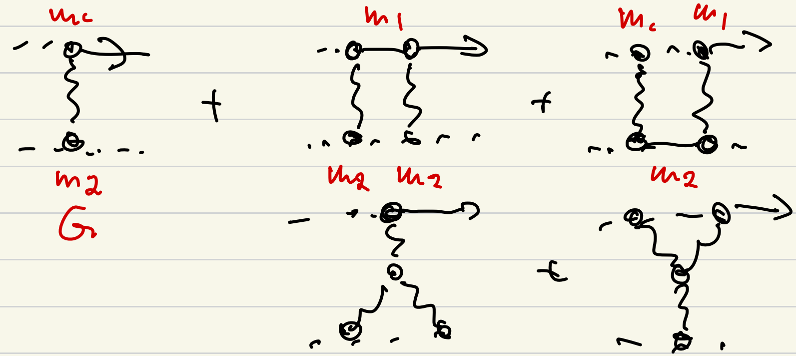
II IMPULSE:

$$\Delta p_1^\mu = m_1 \langle \dot{x}_1^\mu \rangle \Big|_{\tau=-\infty}^{\tau=\infty} = m_1 \int d\tau \, \langle \ddot{x}_1(\tau) \rangle$$

$\parallel$

$\ddot{z}_1$ kl

$$\underset{\text{FT.}}{=} - m_1 \, \omega^2 \left\langle Z_1^\mu(\omega) \right\rangle \Big|_{\omega \to 0}$$

$\Delta p_1^\mu =$  $\omega = 0$ $=$ 

G $\qquad + \qquad G^2$

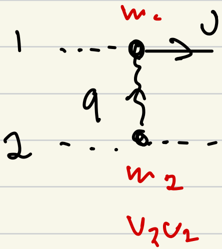
$$+ \; G^3 \; + \; G^4 \; + \; G^5 \; \text{(STATE OF THE ART)} \; + \; \ldots$$

# DIAGS.	26	201	426 + (613)

SELF FORCE EXPANSION: $\qquad m_1, m_2 \qquad m_1 \gg m_2$

IMPULSE @ 1PM

$$b = b_2 - b_1$$



$$= 8\pi i\, G m_1 m_2 \int_q e^{iq\cdot b}\, \delta(v_2\cdot q)\,\delta(q\cdot v_1)\, \frac{P_{\mu\nu; v_2 v_2}\, q^\mu}{q^2}$$

$$P_{\nu q; v_2 v_2} = q^\nu (2\gamma^2 - 1)$$

$$\gamma = v_1 \cdot v_2 = \frac{1}{\sqrt{1-\vec{v}^2/c^2}}$$

$$\Delta p_1^\mu = 4\pi G m_1 m_2 (2\gamma^2-1)\, \partial_{b_1^\mu} \underbrace{\int_q \frac{\delta(q\cdot v_1)\,\delta(q\cdot v_2)}{q^2}\, e^{iq\cdot b}}_{= c + \frac{1}{2\pi\sqrt{\gamma^2-1}} \log|b|}$$

$$= 2 G m_1 m_2 \frac{2\gamma^2-1}{\sqrt{\gamma^2-1}}\, \frac{b^\mu}{|b|^2}$$

SCATTERING ANGLE: $|\Delta p_1^\mu| = 2 p_\infty \sin(\theta/2)$

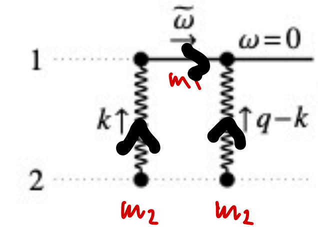
$$p_\infty = m_1 m_2 \sqrt{\gamma^2 - 1} / E$$

$\Rightarrow$
$$\boxed{\frac{\theta}{\Gamma} = \frac{GM}{|b|} \frac{2(2\gamma^2 - 1)}{\gamma^2 - 1}}$$

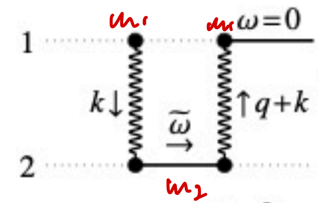
$$\Gamma = \sqrt{1 + 2\nu(\gamma - 1)} \qquad \nu = \frac{m_1 m_2}{m_1 + m_2}$$

$$M = m_1 + m_2$$

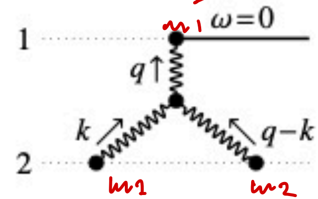
2PM IMPULSE



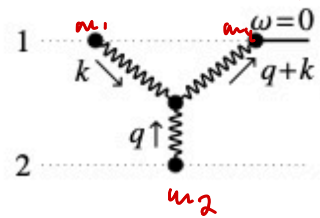
$$= i16\pi^2 G^2 m_1 m_2^2 \int_{k,q} \frac{(q^\mu - k^\mu)\delta(q\cdot v_1)\delta(q\cdot v_2)\delta(k\cdot v_2)}{k^2(k-q)^2(k\cdot v_1 + i0^+)^2} e^{iq\cdot b} \times \left[8\gamma^2(k\cdot v_1)^2 - (2\gamma^2-1)^2 k\cdot(q-k)\right],$$



$$= i16\pi^2 G^2 m_1^2 m_2 \int_{k,q} \frac{(q^\mu + k^\mu)\delta(q\cdot v_1)\delta(q\cdot v_2)\delta(k\cdot v_1)}{k^2(k+q)^2(k\cdot v_2 + i0^+)^2} e^{iq\cdot b} \times \left[8\gamma^2(k\cdot v_2)^2 - (2\gamma^2-1)^2 k\cdot(k+q)\right].$$

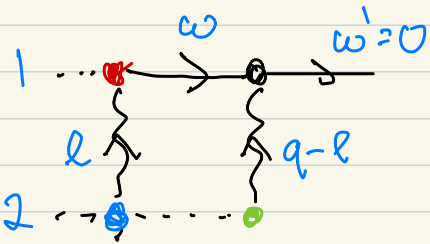


$$= -i64\pi^2 G^2 m_1 m_2^2 \int_{k,q} \frac{q^\mu\delta(q\cdot v_1)\delta(q\cdot v_2)\delta(k\cdot v_2)}{k^2 q^2(k-q)^2} e^{iq\cdot b} \times \left[\gamma^2 q^2 + (k\cdot v_1)^2 + (2\gamma^2-1) k\cdot(k-q)\right],$$



$$= i64\pi^2 G^2 m_1^2 m_2 \int_{k,q} \frac{(k^\mu + q^\mu)\delta(q\cdot v_1)\delta(q\cdot v_2)\delta(k\cdot v_2)}{k^2 q^2(k+q)^2} e^{iq\cdot b} \times \left[\gamma^2 q^2 + (k\cdot v_2)^2 - (2\gamma^2-1) k\cdot(k+q)\right].$$

EXAMPLE:



$\delta(q\cdot v_2)$ $\qquad$ $\omega = \ell\, v_1$ $\qquad$ $\delta(q\cdot v_1)$

$$\int_{q,\ell,\omega} e^{iq\cdot b}\, \frac{\delta(\ell\cdot v_2)\;\overbrace{\delta[(q-\ell)\cdot v_2]}^{=\,\delta(q\cdot v_2)}\;\delta(\ell\cdot v_1-\omega)\;\overbrace{\delta(\omega-(q-\ell)\cdot v_1)}^{=\,\delta(q\cdot v_1)}}{(\omega+i0^+)^2\;\ell^2\;(q-\ell)^2}$$

$v_1^2 = 1 = v_2^2$

$\gamma = v_1\cdot v_2$

$$= \int_q e^{iq\cdot b}\,\delta(q\cdot v_1)\,\delta(q\cdot v_2)\;\underbrace{\int_\ell \frac{\delta(\ell\cdot v_2)}{\ell^2\,(q-\ell)^2\,[\ell\cdot v_1+i0^+]^2}}_{|q|^\alpha\, f[\gamma,\,\epsilon]}$$

@ n-PM ORDER WE HAVE THE GENERAL STRUCTURE

$$\mathcal{J}_{\vec{n}} = \int_q \delta(q\cdot v_1)\,\delta(q\cdot v_2)\, e^{-iq\cdot b} \int_{\ell_1 \dots \ell_{n-1}} \frac{\text{num}}{D_1^{n_1}\dots D_j^{n_j}} \underbrace{\delta(\ell_1\cdot V_*) \dots \delta(\ell_{n-1}\cdot V_*)}$$

$V_* = \begin{cases} v_1 \\ v_2 \end{cases}$ MIRRORS $m_1^{\#}$ $m_2^{\#}$

WITH

$$D_i = \begin{cases} (r^0 + i0^+)^2 - \vec{r}^{\,2} \\ \\ r\cdot V_* + i0^+ \end{cases} \qquad r^\mu = \sum_{a=1}^{n-1} c_a \ell_a^\mu + d\, q^\mu$$

$\gamma = v_1 \cdot v_2$

$$\tilde{I}_{n_1,\dots,n_j}[1] = \int_{\ell_1\dots\ell_{n-1}} \frac{1}{D_1^{n_1}\dots D_j^{n_j}}\, \delta(\ell_1\cdot V_*)\dots\delta(\ell_{n-1}\cdot V_*) = \frac{1}{|q|^\alpha}\, I_{n_1,\dots,n_j}[\gamma,\varepsilon]$$

2PM- INTEGRAL

$$I_{n_1 n_2 n_3}[1] = \int_\ell \frac{f(\ell\cdot v_2)\quad [1]}{[\ell^2]^{n_1}\,[(\ell-q)^2]^{n_2}\,[\ell\cdot v_1 + i0^+]^{n_3}} = |q|^{d-1-2n_1-2n_2-n_3}\; I_{\vec n}[\gamma, \epsilon]$$

$$[I] = d-1-2n_1-2n_2-n_3$$

Symmetries:

1) $I_{n_1 n_2 n_3} = I_{n_2 n_1 n_3}$ $\qquad \ell \to \ell + q$

2) $I_{0 n_2 n_3} = 0 = I_{n_1 0 n_3}$ $\qquad$ Scaleless

TENSOR RED.

$$\hat v_1^\mu = \frac{\gamma v_2^\mu - v_1^\mu}{\gamma^2 - 1} \qquad \hat v_i \cdot v_j = \delta_{ij}$$

$$I_{\vec n}[\ell^\mu] = c_1 \hat v_1^\mu + c_2 \hat v_2^\mu + c_q\, q^\mu / |q|^2$$

$$I_{\vec{n}}[\ell\cdot v_2] = 0 = c_2 \qquad I_{n_1 n_2 n_3}[\ell\cdot v_1] = I_{n_1 n_2 n_3 -1}[1] = c_1$$

$$I_{n_1 n_2 n_3}[\ell\cdot q] = c_q \qquad \ell\cdot q = \frac{1}{2}\left[\ell^2 - (\ell-q)^2 + q^2\right]$$

$$\parallel$$

$$\frac{1}{2} I_{n_1-1\, n_2\, n_3}[1] - \frac{1}{2} I_{n_1\, n_2-1\, n_3}[1] + \frac{q^2}{2} I_{n_1 n_2 n_3}[1]$$

HIGHER RANKS:

$$I_{\vec{n}}[\ell^\mu \ell^\nu] = \sum_{ij} c_{ij}\, v_i^{(\mu} v_j^{\nu)} + \sum_i c_i\, v_i^\mu q^\nu + c\, q^\mu q^\nu + d\, \eta^{\mu\nu}$$

$$\Rightarrow \quad I_{112}[\ell\cdot(\ell-q)(q-\ell)^\mu] = -\frac{q^2}{4}\left[I_{112}\, q^\mu - 2 I_{111}\, v_1^\mu\right]$$

SCALAR INTEGRAL @ 1-LOOP CAN BE PERFORMED WITH STD. QFT TECHNIQUES:

$$I_{n_1 n_2 n_3} = |q|^{d-1-2n_1-2n_2-n_3} (-i)^{2n_1+2n_2+n_3} \frac{2^{n_1} (4\pi)^{\frac{1-d}{2}}}{[\gamma^2-1]^{n_3/2}}$$

$$\Gamma_{n_1 n_2 n_3}[3-2\epsilon]$$

FINAL F.T.

$$\int_q e^{iq\cdot b} \delta(q\cdot u_1)\,\delta(q\cdot u_2)\,|q|^n = \frac{1}{\sqrt{\gamma^2-1}} \frac{2^n}{\pi^{(d-2)/2}} \frac{\Gamma\left(\frac{d-2+n}{2}\right)}{\Gamma\left(-n/2\right)}$$

$$\propto \quad |P_{12} \cdot b|^{2-d-n}$$

$$P_{12}^{\mu\nu} = \eta^{\mu\nu} - \sum_{i=1}^{2} v_i^\mu \hat{v}_i^\nu$$

$b = b_2 - b_1$

$$\boxed{\Delta p_1^\mu = \frac{G m_1 m_2 b^\mu}{|b|^2} \left[\frac{2(2\gamma^2-1)}{\sqrt{\gamma^2-1}} + \frac{3\pi}{4} \frac{5\gamma^2-1}{\sqrt{\gamma^2-1}} \frac{G(m_1+m_2)}{|b|} \right] + (m_2 \hat{v}_1^\mu - m_1 \hat{v}_2^\mu)\, 2 G^2 m_1 m_2 \frac{(2\gamma^2-1)^2}{(\gamma^2-1)|b|^2} + \mathcal{O}(G^3)}$$

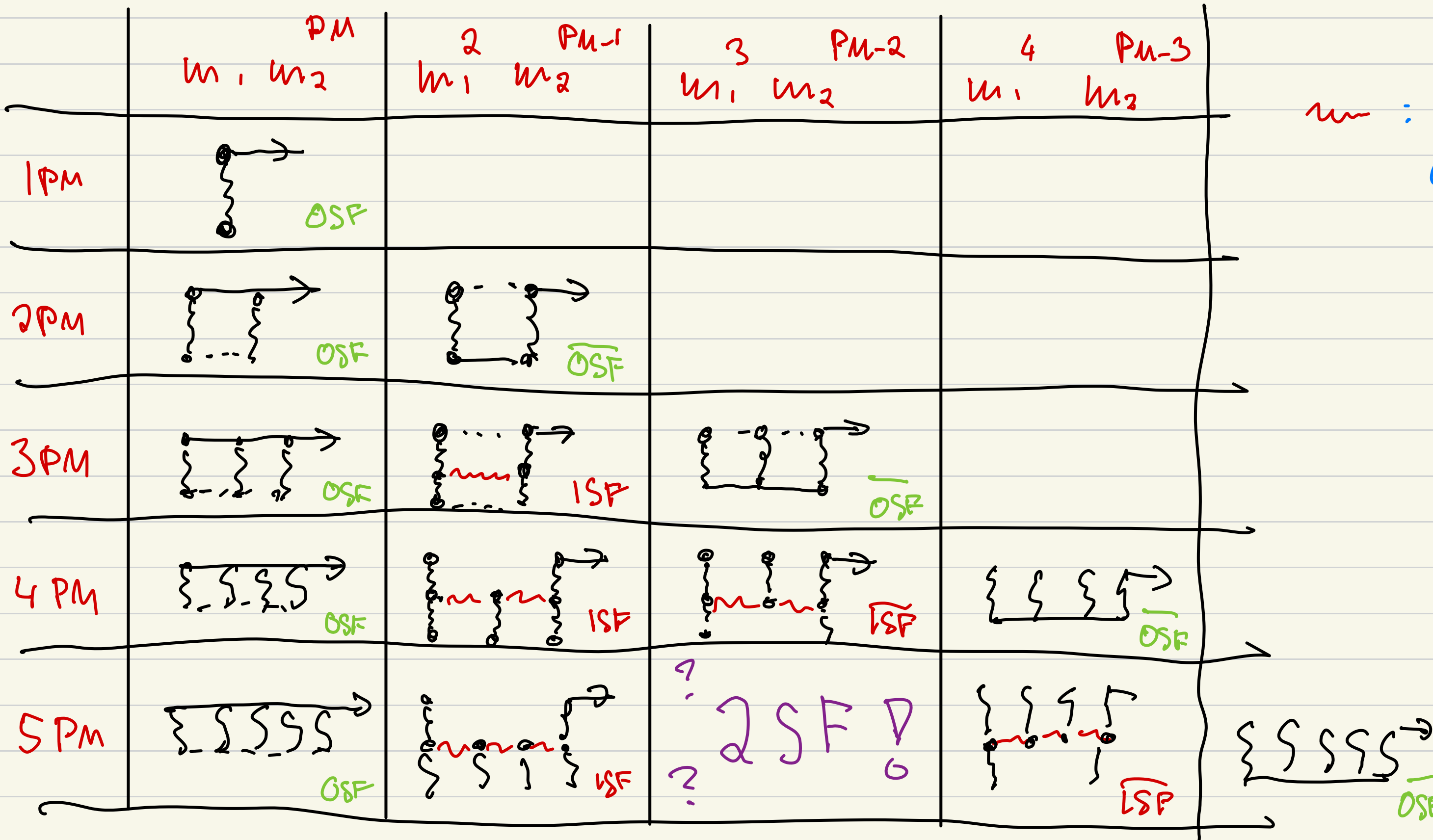
$\Delta p_1 = -\Delta p_2$ upon $1 \leftrightarrow 2$ SWAP

$\Longleftrightarrow$ CONSERVATIVE

$p_1^2 = (p_1 + \Delta p_1)^2 = m^2$

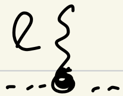
$\rightarrow \quad p_1 \cdot \Delta p_1 = \Delta p_1^2$

HIGHER PM ORDERS : PM VS. SF ORDERS



~ : ACTIVE GRAVITON

$i0^+$ reboot



$$\sim \int_{\ell} \frac{\delta(\ell\cdot v)}{\ell^2} = \int_{\ell} \frac{\delta(\ell^0)}{(\ell^0)^2-\vec{\ell}^{\,2}} = \int_{\vec{\ell}} \frac{1}{-\vec{\ell}^{\,2}} = \text{POTENTIAL}$$

RADIAL GRAVITON $\quad (\ell^0,\vec{\ell}) \sim (v, v)$

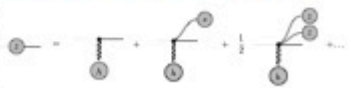
POTENTIAL " $\quad (\ell^0,\vec{\ell}) \sim (v, 1)$

HIGH PRECISION WQFT COMPUTATIONS: WORKFLOW

[Driesse,Jakobsen,Klemm,Mogull,Nega,JP,Sauer,Usovitsch]

Generate graphs & Insert Feynman rules

Berends-Giele recursion



(5PM: 6g vertex)

↓

Reduction & planarization to scalar integral family

$$\int \prod_i d^D l_i \frac{l_{i_1}^{\mu_1} \dots l_{i_n}^{\mu_n}}{\prod_i D(l_i)}$$

↓

$i0^+$ Integration by parts & Master Integrals

Kira 3.0

Reduced symmetries due to retarded prop.

→

Solve masters by Diff. Eq. technique

Single scale integrals

$$x = \gamma - \sqrt{\gamma^2 - 1}$$

$$\gamma = v_1 \cdot v_2$$

↑

$i0^+$ Fix boundary conditions in static limit (PN) $\gamma \to 1$

Method of regions (potential & radiation gravitons)

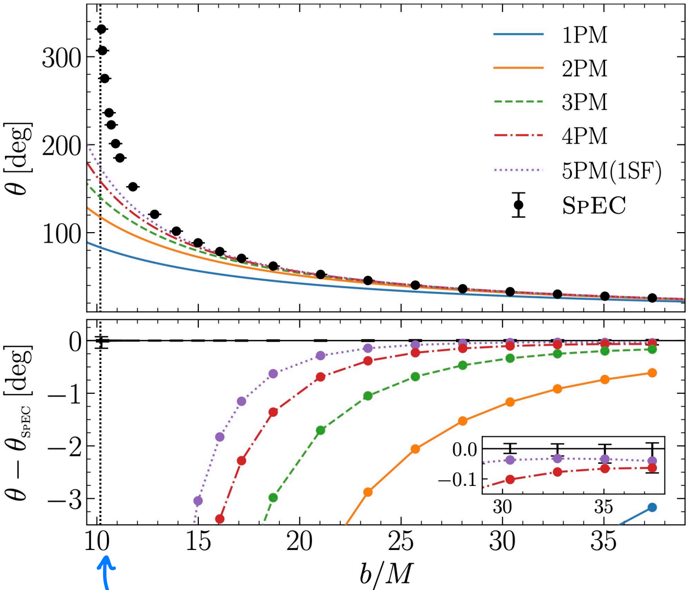
↑

Δp_i^μ $f_{\mu\nu}(u, \hat{x})$ ΔS_i^μ

SCATTERING ANGLE: COMPARISON TO NUMERICAL RELATIVITY

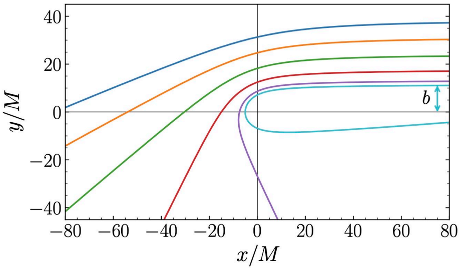
[Long, Pfeiffer, Kidder, Scheel, 2511.10196]

Scattering angle



SEPARATRIX, $b_{\text{crit.}}$

Trajectories:



Parameters:

$$m_1 = m_2$$

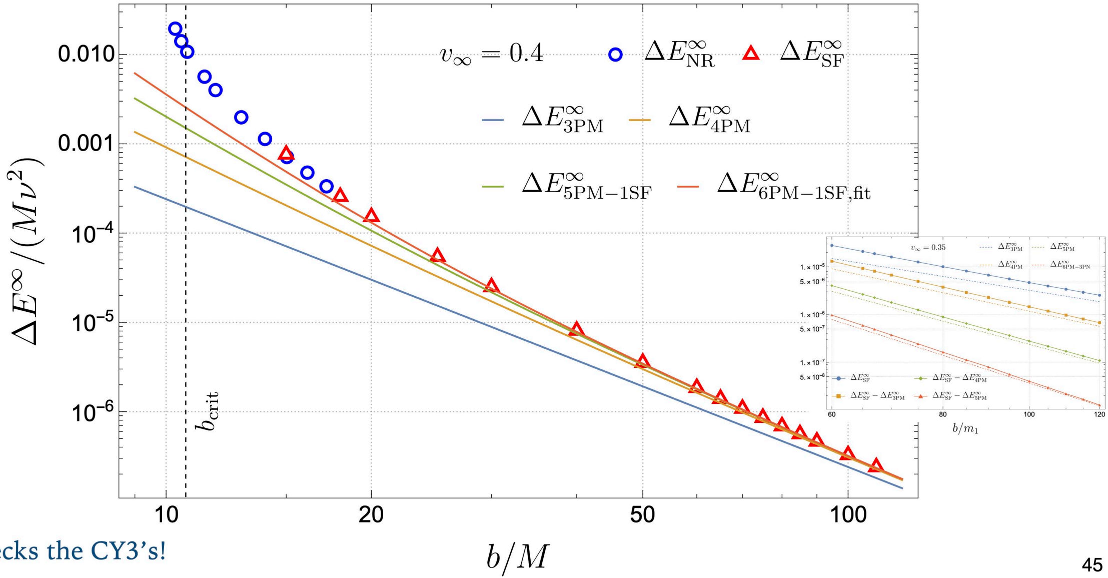
$$v = 0.296c$$

$$M = m_1 + m_2$$

RADIATED ENERGY: COMPARISON TO SEMI-NUMERICAL SELF-FORCE

[Warbuton, 2512.02274]

Radiated energy



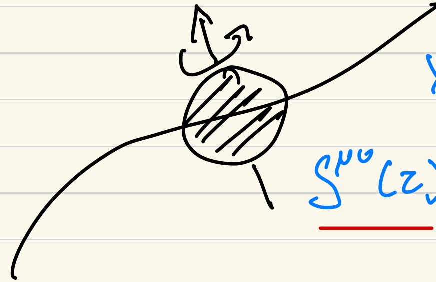
Checks the CY3's!

SCATTERING SPINNING BLACK HOLES: STATE-OF-THE-ART

Higher orders in spin are **suppressed by physical PM counting**: $a_i^\mu = G m_i \chi_i^\mu \,, \qquad |\chi_i| < 1$

	S^0	S^1	S^2	S^3	S^4	S^5
1PM (tree level)	G	G^2	G^3	G^4	Hoogeveen, Jakobsen, JP `25	G^6
2PM (1 loop)	G^2	G^3	G^4	G^5	Guevara, Ochirov, Vines '18 Chen, Chung,Huang, Kim '21 Haddad, Jakobsen, **Mogull** JP '24	Bern, Kosmopoulos, Luna, Roiban, Teng '22; Aoude, Haddad, Helset '23; Bautista '23; Chen, Wang '24; Bohnenblust, Cangemi, Johansson, Pichini '24
3PM (2 loops)	Bern, Cheung, Parra-Martinez, Ruf, Herrmann, Roiban, Shen, Solon, Zeng, Kälin, Liu, Porto, Di Vecchia, Heissenberg, Russo, Veneziano, Travaglini, Brandhuber, Damgaard, Planté, Vanhove, Bjerrum-Bohr	**Jakobsen, Mogull '22** Akpinar, Cordero, Kraus, Ruf, Zeng '24	**Jakobsen, Mogull '22** Akpinar, Cordero, Kraus, Ruf, Zeng '24	Akpinar, Cordero, Kraus, Smirnov, Zeng '25 Haddad, Jakobsen, **Mogull**, JP `25	Akpinar, Cordero, Kraus, Smirnov, Zeng '25 Haddad, Jakobsen, **Mogull**, JP `25	G^8
4PM (3 loops)	Dlapa, Kälin, Liu, Neef, Porto, Damgaard, Hansen, Planté, Vanhove, Bern, Parra-Martinez, Roiban, Ruf, Shen Solon, Zeng	**Jakobsen, Mogull, JP, Sauer, Xu '23**	G^6	G^7	G^8	G^9
5PM (4 loops)	**Driesse, Jakobsen, Klemm, Mogull, Nega, JP, Sauer, Usovitsch '24**	G^6	G^7	G^8	G^9	G^{10}

SPIN ON THE WORLD LINE



$x^\mu(\tau)$

$S^{\mu\nu}(\tau)$: SPIN TENSOR

$$S^{\mu\nu} = -S^{\nu\mu}$$

HAMILTONIAN PERSPECTIVE

(van Holten, 2015)

PHASE SPACE: $\{x^a, P_b, S^{ab}\}$

- FLAT SPACE POISSON STRUCTURE:

$$\boxed{\begin{aligned} &\{x^a, P_b\} = \delta^a_b \\ &\{S^{ab}, S^{cd}\} = \eta^{bd} S^{ac} + 3 \text{ more} \end{aligned}}$$

- CURVED SPACE-TIME:

$$\Pi_\mu := P_\mu + \frac{1}{2}\,\omega_{\mu,ab}\, S^{ab} \qquad S^{\mu\nu} = e^\mu{}_a e^\nu{}_b S^{ab}$$

$$g_{\mu\nu} = e_\mu{}^a e_\nu{}^b \eta_{ab}$$

$$\boxed{\begin{array}{ll} \{X^\mu, \pi_\nu\} = \delta^\mu_\nu & \{S^{\mu\nu}, S^{\rho\kappa}\} = g^{\nu\kappa} S^{\mu\rho} + 3 \text{ MORE} \\ \{S^{\mu\nu}, \pi_\rho\} = 2\,\Gamma^{[\mu}{}_{\rho\kappa} S^{\nu]\kappa} \;\hat{=}\; D_\rho S^{\mu\nu} & \partial_\rho S^{\mu\nu} = 0 \\ \{\pi^\mu, \pi^\nu\} = -\frac{1}{2} R_{\mu\nu;ab} S^{ab} & \end{array}}$$

LAGRANGIAN DESCRIPTION ? TRADITIONAL WAY : Body field $\Lambda^\mu{}_a$

Better solution: Take $S^{\mu\nu}$ as composite (inspired by string theory)

INTRODUCE A COMPLEX VECTOR $\boxed{\alpha^a(\tau)}$, $\bar{\alpha}^a(\tau) = (\alpha^a(\tau))^+$

Postulate:

$$\boxed{\{\alpha^a, \bar{\alpha}^b\} = -\frac{i}{m}\eta^{ab}}$$

Spin-tensor:

$$\boxed{S^{ab} := 2mi\,\bar{\alpha}^{[a}\alpha^{b]}}$$

$$S^{ab}_{-\infty} \sim \bar{\alpha}^{[a}_{-\infty}\alpha^{b]}_{-\infty}$$

SPINNING PARTICLE:

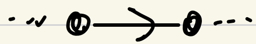
$$\boxed{S_{BH/NS} = -m\int d\tau \Big[\frac{1}{2}g_{\mu\nu}\dot{x}^\mu\dot{x}^\nu + i\,\bar{\alpha}_\mu D_\tau\alpha^\mu + \frac{1}{2}R_{\bar{\alpha}\alpha\bar{\alpha}\alpha} + C_E\, R_{\bar{\alpha}\dot{x}\alpha\dot{x}}\,\bar{\alpha}\cdot\alpha + \mathcal{O}(R^{n\geq 1}, S^{n\geq 3})\Big]}$$

$$D_\tau\alpha^\mu = \dot{\alpha}^\mu + \Gamma^\mu_{\dot{x}\alpha}$$

BACKGROUND FIELD EXP.

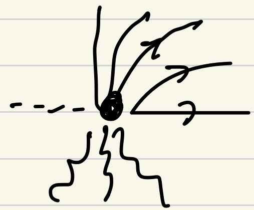
$$\alpha^\mu(\tau) = \alpha^\mu_{-\infty} + \alpha'^\mu(\tau)$$

PROGATOR:



$$\boxed{\langle \alpha^\mu(\omega)\,\bar{\alpha}^\nu(-\omega)\rangle_0 = \frac{-i\eta^{\mu\nu}}{m(\omega+i0^+)}}$$

NEW VERTICES:



$R^1 S^n$ $\quad n \leq 4$

$\mathcal{L}_{(R^1 S^s, n)}$	$s=2$	$s=3$	$s=4$
$n=1$	$\frac{1}{m^2}(S\cdot S)^{\mu\alpha} R_{\mu\dot{z}\alpha\dot{z}}$	$\frac{1}{m^3} S^{\mu\nu}(S\cdot S)^{\sigma\alpha} R_{\mu\nu\alpha\dot{z};\sigma}$	$\frac{1}{m^4}(S\cdot S)^{\rho\sigma}(S\cdot S)^{\mu\alpha} R_{\mu\dot{z}\alpha\dot{z};\rho\sigma}$
$n=2$	$\frac{1}{m^2} S^{\mu\nu} S^{\alpha\beta} R_{\mu\nu\alpha\beta}$	$\frac{1}{m^3} S^{\dot{z}\sigma} S^{\mu\nu} S^{\alpha\beta} R_{\mu\nu\alpha\beta;\sigma}$	$\frac{1}{m^4}(S\cdot S)^{\rho\sigma} S^{\mu\nu} S^{\alpha\beta} R_{\mu\nu\alpha\beta;\rho\sigma}$

$R^2 S^n$ $\quad n \leq 4$

$\mathcal{L}_{(R^2 S^s, n)}$	$s=0$	$s=2$	$s=4$
$n=1$	$\lambda^4 R_{\mu\dot{z}\nu\dot{z}} R^{\mu\dot{z}\nu\dot{z}}$	$\frac{\lambda^2}{m^2} R_{\mu\dot{z}\alpha\dot{z}} R^{\mu}{}_{\dot{z}\beta\dot{z}}(S\cdot S)^{\alpha\beta}$	$\frac{1}{m^4}\left(R_{\mu\dot{z}\nu\dot{z}}(S\cdot S)^{\mu\nu}\right)^2$
$n=2$	$\lambda^4 R_{\mu\nu\alpha\beta} R^{\mu\nu\alpha\beta}$	$\frac{\lambda^2}{m^2} R_{\mu\nu\alpha\dot{z}} R^{\mu\nu}{}_{\beta\dot{z}}(S\cdot S)^{\alpha\beta}$	$\frac{1}{m^4} R_{\mu\nu\sigma\dot{z}} R_{\alpha\beta\rho\dot{z}} S^{\mu\nu} S^{\alpha\beta}(S\cdot S)^{\sigma\rho}$
$n=3$		$\frac{\lambda^2}{m^2} R_{\mu\dot{z}\nu\sigma} R^{\mu\dot{z}\nu\dot{z}}(S\cdot S)^{\dot{z}\sigma}$	$\frac{1}{m^4}\left(R_{\mu\nu\alpha\beta} S^{\mu\nu} S^{\alpha\beta}\right)^2$
$n=4$		$\frac{\lambda^2}{m^2} R_{\mu\nu\alpha\beta} R^{\mu\nu}{}_{\sigma\dot{z}} S^{\dot{z}\sigma} S^{\alpha\beta}$	$\frac{1}{m^4} R_{\mu\dot{z}\nu\dot{z}} R^{\mu}{}_{\alpha\beta\dot{z}}(S\cdot S)^{\dot{z}\nu}(S\cdot S)^{\alpha\beta}$

$\dot{z} = \dot{x}$ $\qquad \lambda$: Length scale.

STRUCTURE OF HIGHER DIM. OPERATORS

$R^n S^s$

$$S_{NM} = -m \int d\tau \sum_A C_A \mathcal{L}_A \qquad A \in (R^n, S^s; n)$$

WILSON-COEF.

DECOMPOSITION OF SPIN-TENSOR:

PAULI-LUBANSKI VECTOR S^μ

SSC-VECTOR: Z^μ

$$\boxed{S^{\mu\nu} = \varepsilon^{\mu\nu\rho\sigma} \dot{x}_\rho S_\sigma + 2 Z^{[\mu} \dot{x}^{\nu]}} \quad (*)$$

$\#$ DoF: $3+3=6$

SPIN-SUPPLEMENTARY CONDITION!

$$\boxed{Z^\mu = 0}$$

CONSTRAINT!

INVERT (+): $S^{\mu} = \frac{1}{2}\varepsilon^{\mu\nu\rho\sigma}\dot{x}_{\nu}S_{\rho\sigma}$ $Z^{\mu} = S^{\mu\nu}\dot{x}_{\nu}$

NEEDS TO HAVE $\boxed{\frac{d}{d\tau}\left[S^{\mu\nu}\dot{x}_{\nu}\right] = 0}$

TRANSLATE SSC $Z^{\mu}=0$ TO $\bar{\alpha}^{\mu}, \alpha^{\nu}$: $\boxed{\frac{d}{d\tau}\left(\alpha\cdot\dot{x}\right) = 0}$

CONSTRAINS THE WILSON COEFF:

$(R^1, S^{2,3,4})$: 3 free coeff. $(R^2, S^{2,3,4})$: 6 free coeff

FOR KERR-BH:

(R^1, S^4) all coeffs fixed !

$(R^2, S^{4>4})$?